







- These sorts of questions have ensured that optics has a long and engaging history.
- Mirrors were known to the ancients, eyeglasses were known by the thirteenth century, and, of course, the telescope was invented by Galileo around 1608.
- The law of refraction was discovered by Willebrord Snell in 1621 and the phenomenon of diffraction was
  observed by both Francesco Maria Grimaldi and Robert Hooke by the mid-1600s.
- Sir Isaac Newton made great contributions to optics, proposing that 'white light' was a combination of all colors, and formulating a particle, or corpuscular, theory of light.
- At roughly the same time (the latter half of the seventeenth century), the Dutch physicist Christiaan Huygens
  proposed a powerful wave theory of light.



• As we shall discover, most of the history of optics is dominated by the debate over the nature of light:

Is light a particle or a wave, or is it something in between (a wavicle?)?

In quantum mechanics, a **wavice** is a <u>wave-particle</u>. An entity which simultaneously has the properties of a wave and a particle.



- Another important figure in the history of optics is Thomas Young, an Englishman who revived the wave theory at the beginning of the nineteenth century by adding to it the principle of superposition.
- The French scientist Augustin Jean Fresnel, also an advocate of the wave theory, proposed a mechanistic description of light on the basis of it being a transverse oscillation through the ether, rather than a longitudinal one as had previously been assumed. The corpuscular theory seemed in very bad shape indeed.
- By 1845 Michael Faraday had performed several experiments showing that the plane of polarization could be altered by magnetic fields. This ultimately led to James Clerk Maxwell's brilliant unification of optics and electromagnetism, when his wave equations predicted that the speed of light should be  $1/\sqrt{\varepsilon_0 \cdot \mu_0}$ , which was remarkably close to the experimental value. Light, then, was an electromagnetic disturbance propagating through the ether.







- As a wave, however, light must have a medium through which to propagate. Towards the end of the nineteenth
  century this medium, called the ether, became increasingly problematic; experiments by Michelson and Morley in
  particular could detect no motion of the ether relative to the earth.
- Such considerations led to Einstein's theory of special relativity and to the discarding of the idea of the ether altogether. Moreover, as the twentieth century progressed, quantum mechanics showed that all particles have a wavelike property; the distinction between waves and particles became less and less clear.
- In this lecture we will treat light usually as a wave, but sometimes as a particle, and as a general rule it is both or either.







#### Photons

- Particles with energy hv
- Important for calculating detector efficiencies, considering scattering processes, beamsplitting, blackbody radiation

#### Waves

- With interference, vibrate charges in a substance have a polarization
- · Important for diffraction limit, seeing, gratings, diffraction spikes, polarization effects in scattering

#### Rays

- Representing the path of photons (or equivalently wavefront normals)
- Important for thinking about optics, image quality, aberrations, etc. (not as fundamental as particle and waves, but
  useful analytically/geometrically)





#### Geometrical vs Physical Optics





#### Geometrical Optics

- Traces "rays" which do not interact with one another
- Purely Geometric (e.g. Snell's law, law of reflection)
  Intersection of individual rays defines a focal plane and gives an image

#### **Physical Optics**

- Wave nature of light
- Wavefronts (rather than rays)
- Wave propagation and interference determine illumination
- Interference, diffraction, etc...







### Chapter

Paris Foulousi

## **Geometrical Optics**

Geometrical optics or ray optics describes light propagation in terms of rays. The ray in geometric optics is an abstraction useful for approximating the paths along which light propagates under certain circumstances.

## **II. Geometrical optics**

FOR REMOTE SENSIN

GEOMETRICAL APPROXIMATION

- Ideal optics = spherical waves from any point in object space are imaged into points in image space
- Corresponding points are called conjugate points
- Focal point = center of converging or diverging spherical wavefront
- Object space and image space are reversible





- Rays are normal to locally flat wave (locations of constant phase)
- Rays are reflected and refracted according to Fresnel equations
- Phase is neglected ⇒ Incoherent sum
- Rays can carry polarization information
- Geometrical optics neglects diffraction effects:  $\lambda \ \rightarrow 0$
- Physical optics  $\lambda > 0$
- · Simplicity of geometrical optics mostly outweighs limitations





# FUNDAMENTALS OF LIGHT

- Light travels in a straight line in constant refractive index medium at speed c/n
- Refractive index n is 1.0 in vacuum, and is related to the permittivity ( $\varepsilon$ ) and permeability ( $\mu$ ) of material

$$\frac{1}{c} = \sqrt{\varepsilon_0 \cdot \mu_0} \qquad \frac{1}{v} = \sqrt{\varepsilon \cdot \mu} \qquad n = \sqrt{\frac{\varepsilon \cdot \mu}{\varepsilon_0 \cdot \mu_0}}$$

Material	Index (n)	notes
air	1.000274	274ppm different from vacuum
water	1.333	Similar for eyeball
quartz	1.458	Aka, fused silica
BK7 glass	1.52	Common optics for lenses
Diamond	2.419	
Silicon	~3.5	CCD material, for instance







#### Refraction

- Light bends at interface between refractive indices (n)
  - Bends more the larger the difference in refractive index
  - Can be effectively viewed as a least time behavior get from A to B faster if you spend less time in the slow medium





#### Optical design considering every surface

• Let's consider a thick piece of glass (n = 1.5) and the light paths associated with it. We define the reflection fraction

$$R = \left[\frac{n_1 - n_2}{n_1 + n_2}\right]^2$$

Using  $n_1 = 1.5$  and  $n_2 = 1.0$  (air)  $\Rightarrow R = \left(\frac{0.5}{2.5}\right)^2 = 0.04 = 4\%$ 







#### Surface shape of perfect lens

LENSES

- Lens material has index of refraction n
- $\overline{oz(r)} \cdot n + \overline{z(r)f} = constant$
- $n \cdot z(r) + \sqrt{r^2 + (f z(r))^2} = constant$
- Solution z(r) is an hyperbola with an eccentricity e = n > 1







OPTICS FOR REMOTE SENSING

LENSES

#### Spherical lenses

- · If two spherical surfaces have same the radius, one can fit them together
- Surface error requirement less than  $\lambda/10$
- Grinding spherical surfaces is easy ⇒ most optical surfaces are spherical
- System miniaturization: Freeform optics involve optical designs with at least one freeform surface which, according
  to the ISO standard 17450-1:2011, has no translational or rotational symmetry about axes normal to the mean
  plane.











#### General lens setup and Real image

- •
- Axis through two centers of curvature is called optical axis
- Surface point on optical axis is called the vertex
- Chief ray through center maintains direction





#### General lens setup and Virtual image

- Object distance S<sub>1</sub> Object height h<sub>1</sub>
- Image distance S<sub>2</sub> Image height h<sub>2</sub>
- Note that the object is closer than the focal length of lens
- Virtual image



## II. Geometrical optics

#### LENSES

#### Thin lens approximation

Thin-lens equation

$$\frac{1}{S_1} + \frac{1}{S_2} = (n-1) \cdot \left(\frac{1}{R_1} - \frac{1}{R_2}\right)$$

 $\frac{1}{S_1} + \frac{1}{S_2} = \frac{1}{f}$ 

Gaussian lens formula

#### Finite imaging

- Rarely image point sources, but extended object
- Object and image size are proportional
- Orientation of object and image are inverted
- (transverse) magnification perpendicular to optical axis

$$M = \frac{h_2}{h_1} = -\frac{S_2}{S_1}$$



#### Thick lenses

Basic thick lens equation

$$\frac{1}{f} = (n-1) \cdot \left(\frac{1}{R_1} - \frac{1}{R_2} + \frac{(n-1) \cdot d}{n \cdot R_1 \cdot R_2}\right)$$

- Reduces to thin lens for  $d \ll R_1 \cdot R_2$ 



# II. Geometrical optics

#### **Principal planes**

- Focal lengths and distances are measured from the principal planes
- Refraction can be considered to only happen at principle planes
- Principle planes only depend on lens properties itself
- Thin lens equation neglects finite distance between principle planes
- Distance between vertices and principal planes is given by











MIRRORS

#### Mirrors vs Lenses

- Mirrors are completely achromatic
- Reflective over very large wavelength range (UV to radio)
- Can be supported from the back
- Can be segmented
- Wavefront error is twice that of surface, lens is (n-1) times surface
- Only one surface to 'play' with



IPSA





- parallel incident rays of light.
- One of the easiest shapes to analyze is the spherical mirror. Such a mirror is not a complete sphere, but a spherical cap a piece sliced from a larger imaginary sphere with a single cut.
- Easy to manufacture
- · Focuses light from center of curvature onto itself
- Focal length is half of curvature
- Tip-tilt misalignment does not matter
- Has no optical axis
- Does not image light from infinity correctly: spherical aberration





f = R/2

MIRRORS

#### Parabolic mirrors

- A parabolic (or paraboloid or paraboloidal) reflector (or dish or mirror) is a reflective surface used to collect or
  project energy such as light, sound, or radio waves.
- Its shape is part of a circular paraboloid (surface generated by a parabola revolving around its axis). The parabolic
  reflector transforms an incoming plane wave traveling along the axis into a spherical wave converging toward the
  focus.
- · Want to make flat wavefront into spherical wavefront
- Distance  $\overline{az(r)} + \overline{z(r) \cdot f} = constant$  and  $z(r) = r^2/2 \cdot R$
- · Perfect image of objects at infinity
- Has clear optical axis









#### **Elliptical mirrors**

- The elliptical mirrors have two focal points: light that passes through one of these before striking the mirror is
  reflected such that it passes through the other.
- Have two foci at finite distances
- Perfectly reimage one focal point into another









#### OPTICAL SYSTEMS

#### Overview

- Combinations of several optical elements (lenses, mirrors, stops)
- Examples: camera "lens", microscope, telescopes, instruments
- Thin-lens combinations can be treated analytically
- Effective focal length:

$$\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2}$$





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#### Overview

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#### Simple thin-lens combinations

- Distance > sum of focal lengths ⇒ real image between lenses •
- Apply single-lens equation successively









#### F-number and numerical aperture

- In optics, an aperture is a hole or an opening through which light travels.
- More specifically, the aperture and focal length of an optical system determine the cone angle of a bundle of rays that come to a focus in the image plane.
- · All optical systems have a place where 'aperture' is limited
- Main mirror of telescopes
- Aperture stop in photographic lenses
- · Aperture typically has a maximum diameter
- · Aperture size is important for diffraction effects





## II. Geometrical optics



OPTICAL SYSTEMS

f/2:

#### F-number and numerical aperture

- The f-number of an optical system (such as a camera lens) is the ratio of the system's focal length to the diameter of the entrance pupil.
- It is a dimensionless number that is a quantitative measure of lens speed, and an important concept in photography. It is also known as the focal ratio, f-ratio, or f-stop.
- It is the reciprocal of the relative aperture. The f-number is commonly indicated using a hooked f with the format f/D, where D is the f-number.

f/4:

- · Describes the light-gathering ability of the lens
- The **f-number** given by F = f/D
- Also called focal ratio or f-ratio, written as: f/F
- The bigger F, the better the paraxial approximation works
- Fast system for F < 2, slow system for F > 2







#### F-number and numerical aperture

- In optics, the numerical aperture (NA) of an optical system is a dimensionless number that characterizes the range of angles over which the system can accept or emit light.
- By incorporating index of refraction in its definition, NA has the property that it is constant for a beam as it goes
  from one material to another, provided there is no refractive power at the interface.
- Numerical aperture definition

 $NA = n \cdot \sin(\theta)$ 

- *n* index of refraction of working medium
- heta is the half-angle of maximum cone of light that can enter or exit the lens
- This notion is important for microscope objectives (n often not 1)





• The numerical aperture in Fibers is the acceptance cone of the fiber determined by the materials





### FOR REMOTE SENSIN **II. Geometrical optics**

IMAGES & PUPILS

#### Image definition

- Every object point comes to a focus in an image plane
- Light in one image point comes from pupil positions
- Object information is encoded in position, not in angle •

#### Pupil definition

- All object rays are smeared out over complete aperture •
- Light in one pupil point comes from different object positions
- Object information is encoded in angle, not in position







#### Aperture and field stops

- Aperture stop limits the amount of light reaching the image
- Aperture stop determines light-gathering ability of optical system
- Field stop limits the image size or angle





IMAGES & PUPILS

#### Spherical aberration of spherical lens

- Different focal lengths of paraxial and marginal rays
- Longitudinal spherical aberration along optical axis
  Transverse (or lateral) spherical aberration in image plane
- Much more pronounced for short focal ratios
- Foci from paraxial beams are further away than marginal rays
- Spot diagram shows central area with fainter disk around it



## II. Geometrical optics

IMAGES & PUPILS

#### Spherical aberration of spherical lens

• Minimizing spherical aberrations?





#### Other possible aberrations

- In optics (especially telescopes), the coma or chromatic aberration in an optical system refers to aberration inherent to certain optical designs or due to imperfection in the lens or other components that results in off-axis point sources such as stars appearing distorted, appearing to have a tail (coma) like a comet.
- The astigmatism is due to Tilted Glass Plate in converging beam. It focuses in two orthogonal directions, but not in both at the same time.





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- The astigmatism is due to Tilted Glass Plate in converging beam. It focuses in two orthogonal directions, but not in both at the same time.
- Petzval or field curvature: image lies on curved surface
- Petzval field flattening: Petzval curvature only depends on index of refraction and focal length of lenses. It is independent of lens position!





#### Distorsion

- In geometric optics, distortion is a deviation from rectilinear projection: a projection in which straight lines in a scene remain straight in an image. It is a form of optical aberration.
- Image is sharp but geometrically distorted
   (

(a) object(b) positive (or pincushion) distortion(c) negative (or barrel) distortion

